OH Observations with the GBT Position Switched, Single-Pointing Observations

More detailed information can be found in the GBT Observer's Guide:

http://www.gb.nrao.edu/scienceDocs/GBTog.pdf

```
# Observing script for spectral line observations of OH using position
switching.
# Reset configuration from prior observation.
ResetConfig()
# Import catalog of flux calibrators and user defined sources.
Catalog(fluxcal)
Catalog('/home/astro-util/projects/quick guide/catalogs/ps OH.cat')
# Define configuration parameters
ps OH config='''
receiver = 'Rcvr1_2'
obstype = 'Spectroscopy'
backend = 'VEGAS'
restfreq = 1665
bandwidth = 23.44
nchan = 32768
vegas.subband= 3276swmode= 'tp'swtype= Noneswfreq= 0, 0swper= 1.0
swper
                   = 1.0
                   = 6.0
tint
                = 0.0
= 'lsrk'
= 'Optical'
= 'lo'
vframe
vdef
noisecal
pol
                   = 'Linear'
notchfilter = 'In'
111
# Configure telescope.
Configure (ps_OH config)
# Slew to your source or calibrator.
Slew('0228+6721')
# Perform position and focus correction on nearby calibrator.
# Leaving the `()' blank will have the system choose your calibrator for
AutoPeakFocus ('0228+6721')
# Slew to your source.
Slew('W3OH')
# Reconfigure after calibrator corrections.
Configure (ps OH config)
# Balance the IF system
Balance()
# OffOn produces two scans each of the specified duration (in seconds)
which tell the GBT to take data for 10 minutes.
OffOn('W3OH', Offset('J2000', '-00:05:00', 0.0, cosv=True), 300)
```

Catalogs To find out more about catalogs: <u>GBT Observer's Guide: Section 6.3</u>

Here is an example of a RA/Dec coordinate system catalog with velocity:

Note: Vel is source velocity in units of km/s. Reference frames can be set using the VDEF keyword. The default is the reference frame previously set. You can also include any number of user defined keywords.

See Observer's guide for more information.

Configurations To find out more about configurations: <u>GBT Observer's Guide: Section 6.2</u>

Here is an example of a position switched configuration for OH observations:

```
# Configuration parameters for spectral line observations of OH using position switching.
ps OH config= '''
receiver = 'Rcvr1_2'
                            # Specifies L-Band receiver for OH
obstype = 'Spectroscopy'
                            # Specifies spectral line observations
backend = 'VEGAS'
                            # Specifies spectral line backend
                            # Specifies rest frequency for OH (MHz)
restfreq = 1665
deltafreq = 0.0
                            # Specifies offsets for each spectral window (MHz)
bandwidth = 23.44
                            # Defined by chosen VEGAS mode (MHz)
                            # Specifies number of channels in spectral window
nchan = 32768
vegas.subband = 1
                            # Specifies single or multiple spectral windows (1 or 8)
                            # Specifies switching mode, switching power with noise diode
swmode = 'tp'
swtype = None
                            # Specifies frequency switching
                            # Specifies length of full switching cycle (seconds)
swper = 1.0
                            # Specifies frequency offset (MHz)
swfreq = 0, 0
                           # Specifies integration time (sec; integer multiple of swper)
tint = 6.0
vframe = 'lsrk'
                           # Specifies velocity reference frame
                           # Specifies Doppler-shifted velocity frame
vdef = 'Optical'
                           # Specifies level of the noise diode, use 'lo' for 'fsw'
noisecal = 'lo'
                            # Specifies 'Linear' or 'Circular' polarization
pol = 'Linear'
                            # Specify 'In' to block 1200-1310 MHz RFI signal
notchfilter = 'In'
111
```

NOTE: Your parameters may differ based on your specific science goals.

Scripts (Scheduling Blocks)

To find out more about scripts: <u>GBT Observer's Guide: Section 6.1</u>

Astrid is used to submit scripts, or Scheduling Blocks, for GBT observations. Astrid is Python based and can incorporate custom user scripts. Here is an example of a basic position switched, tracking observation for OH observing.

```
# Observing script for spectral line observations of OH using position switching.
# Reset configuration from prior observation.
ResetConfig()
# Import catalog of flux calibrators and user defined sources.
Catalog(fluxcal)
Catalog( '/home/astro-util/projects/quick_guide/catalogs/ps_OH.cat')
# Define configuration parameters
ps_OH_config='''
receiver = 'Rcvr1 2'
obstype = 'Spectroscopy'
backend = 'VEGAS'
restfreq = 1665
bandwidth = 23.44
nchan = 32768
vegas.subband = 1
swmode = 'tp'
swtype = None
swfreq = 0, 0
swper = 1.0
tint = 6.0
vframe = 'lsrk'
vdef = 'Optical'
noisecal = 'lo'
pol = 'Linear'
notchfilter = 'In'
111
# Configure telescope.
Configure(ps OH config)
# Slew to your source or calibrator.
Slew('Calibrator')
# Perform position and focus correction on nearby calibrator.
AutoPeakFocus('Calibrator')
# Slew to your source.
Slew('W3OH')
# Reconfigure after calibrator corrections.
Configure(ps_OH_config)
# Balance the IF system.
Balance()
# OffOn produces two scans each of the specified duration (in seconds) which tell
the GBT to take data for 10 minutes, then 1 minute. Repeat Slew, Balance, OffOn commands
for desired sources.
OffOn('W3OH', Offset('J2000', '-00:05:00', 0.0, cosv=True), 300)
OffOn('W3OH', Offset('J2000', '-00:05:00', 0.0, cosv=True), 300)
OffOn('W3OH', Offset('J2000', '-00:05:00', 0.0, cosv=True), 30)
OffOn('W3OH', Offset('J2000', '-00:05:00', 0.0, cosv=True), 30)
```

Data Reduction To find out more about data reduction: <u>GBTIDL User's Guide</u>

Our current data reduction routines are written in IDL. Users can build custom scripts incorporating generic IDL commands. We will run through some common GBT IDL commands below.

From a Green Bank Observatory data reduction machine (Fourier, Arcturus, Planck, Newton, Euclid), log into GBTIDL by typing "gbtidl" from a terminal.

To access the test data presented in this reference guide type 'offline' followed by the project name:

GBTIDL -> offline, 'TGBT21A_504_01'

Note: 'Connecting to file' tells you where the raw data files are located. File updated shows how long ago the last scan was updated.

Note: To view data from a different observing project, replace the (TGBT) with the information for your project:

o Semester number (e.g., AGBT20A) o Project number (e.g., 108) o Session number (e.g., 01)

Note: To access current observations, or see real-time data during an observing session, type 'online' from the command line. The project code is not needed in online mode.

Type 'summary' to view your observations:

GBTIDL Scan	-> summa Source		Proc	Se	q RestF	[:] n	IF	nInt	nFd /	Az El
14	W30H	0.0	OffOn		1.665	1	11	1	323.4	41.3
15	W30H	0.0	0ff0n	_	1.665	1	11	1	323.3	42.4
16	W30H	0.0	0ff0n	1	1.665	1	11	1	323.4	41.0
17	W30H	0.0	0ff0n	2	1.665	1	11	1	323.3	42.1
18	W30H	0.0	0ff0n	1	1.665	1	11	1	323.5	40.7
19	W30H	0.0	0ff0n	2	1.665	1	11	1	323.3	41.8
26	W30H	0.0	0ff0n	1	1.665	1	6	1	323.7	39.6
27	W30H	0.0	0ff0n	2	1.665	1	6	1	323.5	40.7
28	W30H	0.0	0ff0n	1	1.665	1	6	1	323.8	39.4
29	W30H	0.0	0ff0n	2	1.665	1	6	1	323.5	40.5
30	W30H	0.0	0ff0n	1	1.665	1	6	1	323.8	39.2
31	W30H	0.0	OffOn	2	1.665	1	6	1	323.6	40.3

Note: For more information on what each column is, please see the GBTIDL User's Guide <u>GBTIDL</u> <u>User's Guide: Section 4.7</u>.

Use the 'getps' command to view your position switched observations. Include which scan you would like to look at. Here, we will analyze scan 29.

```
GBTIDL -> getps, 29
```

File Options LeftClick	GHz LSR OPTICAL	Abs Uncome Auto Update	Print	
Scan 29 2021—03—16 Brenne Gregory	V : 0.0 OPTI-L3 Int : 00 00 14.6 LST : +07 25 03.7	SR F0 : 1.66540 Fsky : 1.66538 BW : 23.4375 M	GHz IF: O	Tsys: 20.34 Tcal: 1.46 04_01 OffOn
02 27 03.55 +61 52 2	21.9	W3OH	Az: 323.5	EI: 40.5 HA: 4.97
10000 100000 100000 100000 10000 10000 10000 10000 10000 100000				
1.655	1.660	1.665 LSR Frequency (GHz)	1.670	1.675 Thu Mar 18 10:53:32 2021
Left Click: Null X: 1,656768 Y	: -476,7500 Zoom Level: 0	Auto Update: Off		

Note: This will give you the x-axis in frequency and one polarization. If polarization is not specified, the default is 0.

Note: You can then change the x-axis to velocity, using the buttons on the GUI interface. You can do this using the fourth button on the GUI display.

To get the second polarisation, you can type:

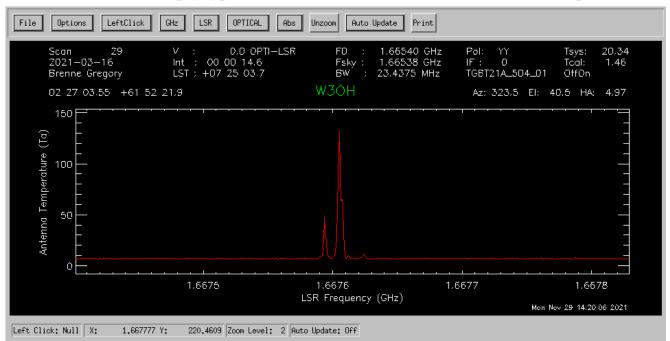
GBTIDL -> getps, 28, plnum=1

To stack/average multiple scans together to improve signal to noise in the spectrum:

GBTIDL -> getps, 28 GBTIDL -> accum GBTIDL -> getps, 30 GBTIDL -> accum GBTIDL -> ave

To zoom in, you can use use the middle mouse button, clicking twice to specify the corners of a zoom box.

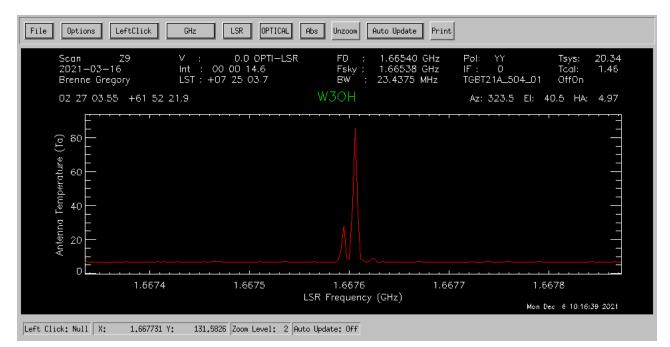
You may also use the 'setxy' command, placing a stretchable box on the plot. In order to unzoom, simply type 'unzoom' or click the *Unzoom* button on the plotter.



Here we have zoomed in to the 1667 MHz OH Line. You can do all this for all of your sources.

To smooth out your spectra, you can use the 'gsmooth' command:

```
GBTIDL -> getps, 28
GBTIDL -> gsmooth, 4, /decimate
```



Saving and/or exporting your data can be done in multiple ways. All of these procedures are located in the <u>GBTIDL User's Guide: Section 9.</u>

One example of this is the write_ascii command:

GBTIDL -> write_ascii, 'data'
ASCII file written:data

This will output the data into the currently directory.